

Langmuir probes and accelerometers for MMOD Observations



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A few means to observe MMOD (≤ 10 cm)



- **MMOD (several mm to several cm) observations by solitons**
 - Coherent Density fluctuations
 - Magnetic field fluctuations
 - Electric field fluctuations
- **MMOD ($<$ several mm) observations by impact on spacecraft**
 - plasma discharges
 - Impact vibrations/sudden change in spacecraft velocity

Solitons and Wakes [MMOD (< several cm)]



- (Solitons – upstream) and (Wakes – downstream)
- At High LEO altitudes (>700 km)...

	Wakes*	Solitons**
Wave amplitude in Density	~+/-2x (1e3/CC) Turbulent structure	~+/-2x (1e3/CC) Coherent structure
Scale size	Multiple tens of cms	Tens of cms
Electric field	???	mV/cm (?)
Magnetic field	???	Few nT (?)

- Langmuir probes requirements
 - Measure fluctuations at 1e2/CC scale (bit better than needed!)
 - For spacecrafts moving at 8 km/sec, sample at ~80 KHz
 - Deploy probe(s) outside the self-wake to detect the soliton
 - Multiple deployed probes (may) help determine direction
 - Measure spacecraft charging to better than 1 mV

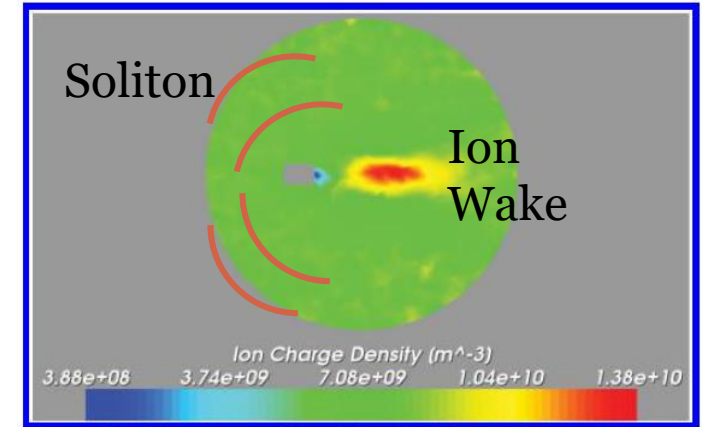


Fig. 5 Ion charge density through the principal axis along the spacecraft velocity vector.

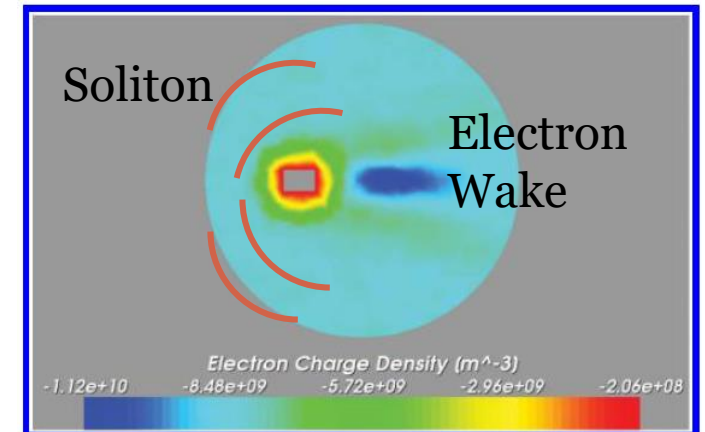
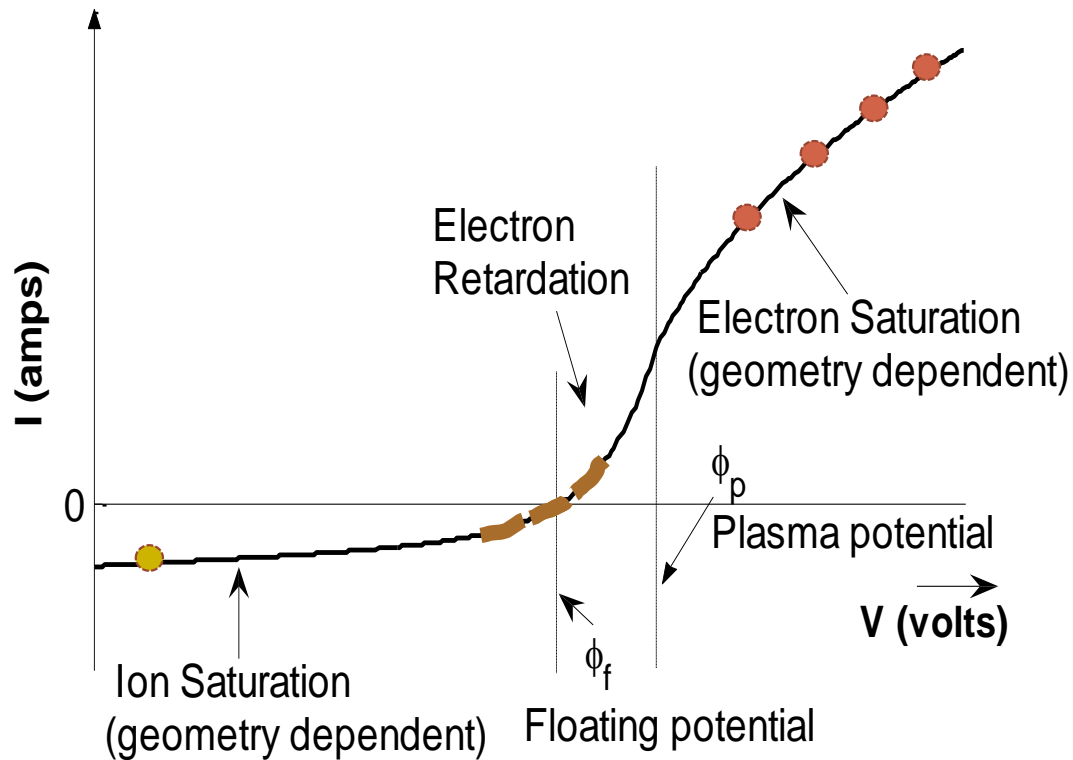


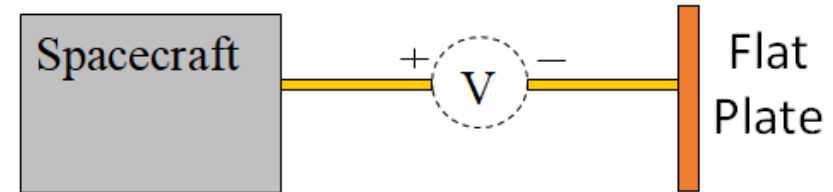
Fig. 6 Electron charge density through the principal axis along the spacecraft velocity vector.

SAIL Langmuir Probes

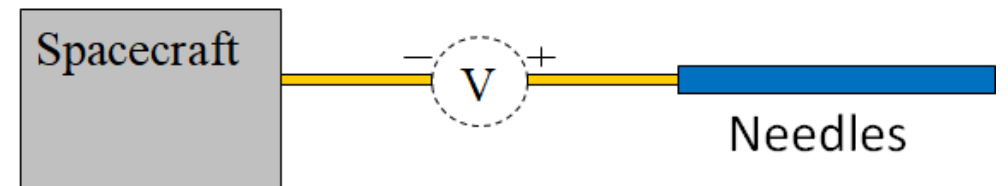
- Sweeping probes, fixed bias probes, floating probes and all geometries



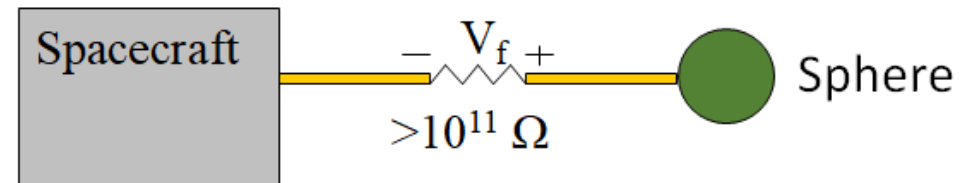
Negative Fixed-bias Langmuir Probe (PIP)



Positive Fixed-bias Langmuir Probe (mNLP)



Floating Potential Probe (FPP)



SAIL Langmuir Probes ... performance

- Density : $1e2/CC$ to $1e7/CC$
- Sampling frequency: Current max ~ 90 KHz
- Needle dimensions: 1 mm diameter – 5 cm length
- Deployable on long booms
- Multiple probes can give directional information!
- 1.6m separation between two-needles takes 0.1 msec to travel at 16 km/s. Resolvable at 80 Khz sampling

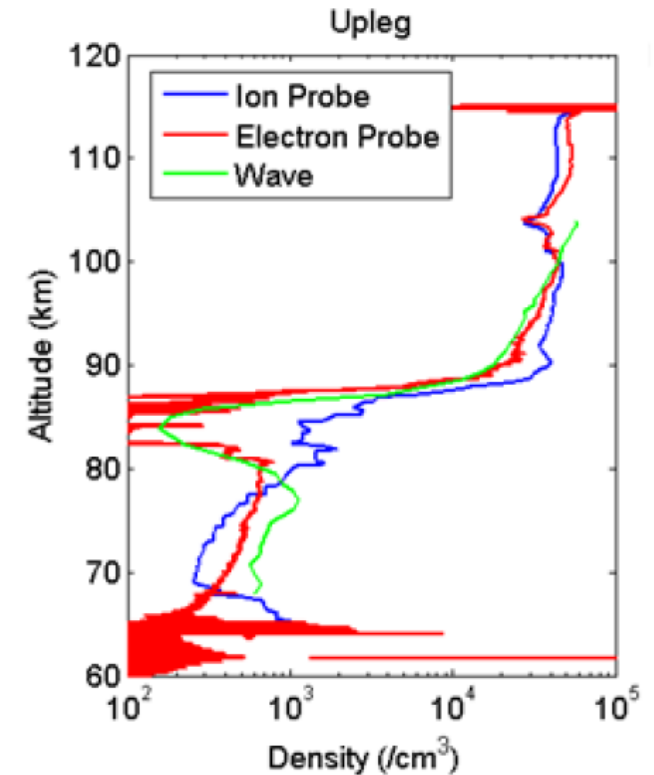
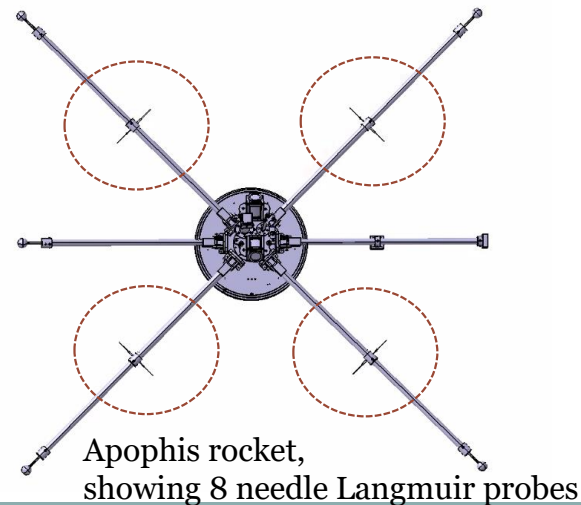
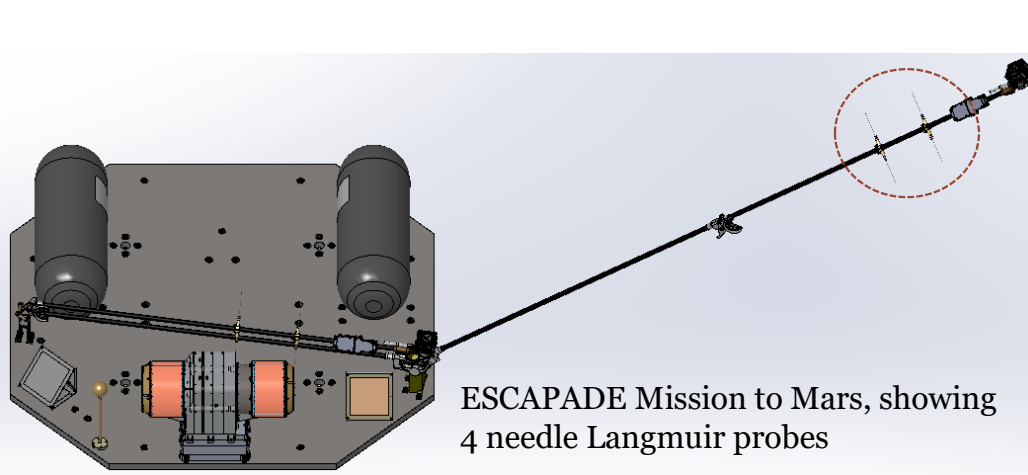


Fig.12 WADIS 1 data showing an electron bite-out, that the Ion probe is immune to.

Magnetic Field



- ~ 1 nT B Fields have been found associated with solitons in Ulysses spacecraft data.
- It is in solar wind, so not clear if that impacts HLEO solitons
- In any case, SAIL miniature magnetometer meticulously calibrated to give 1 nT resolution at 1 Hz... and 3-5 nT at 10 Hz.

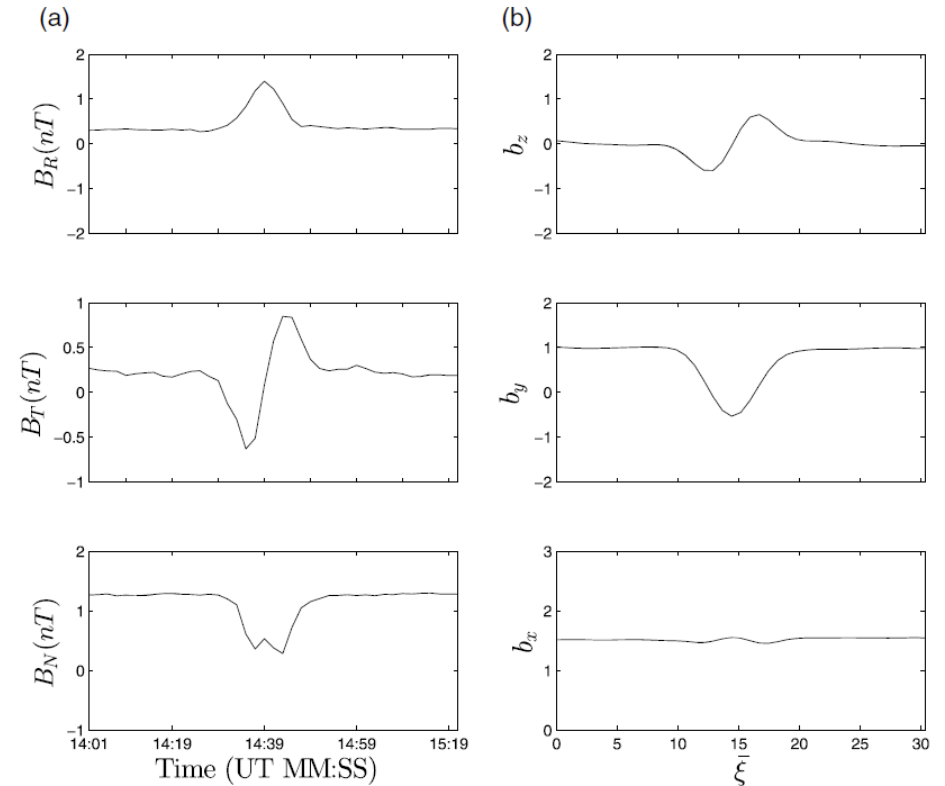


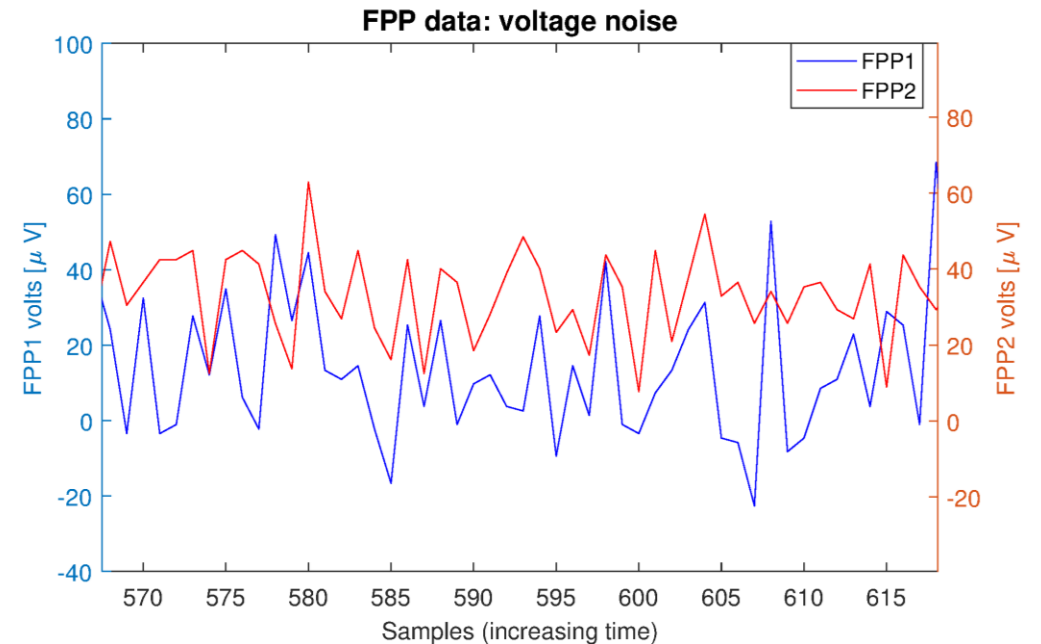
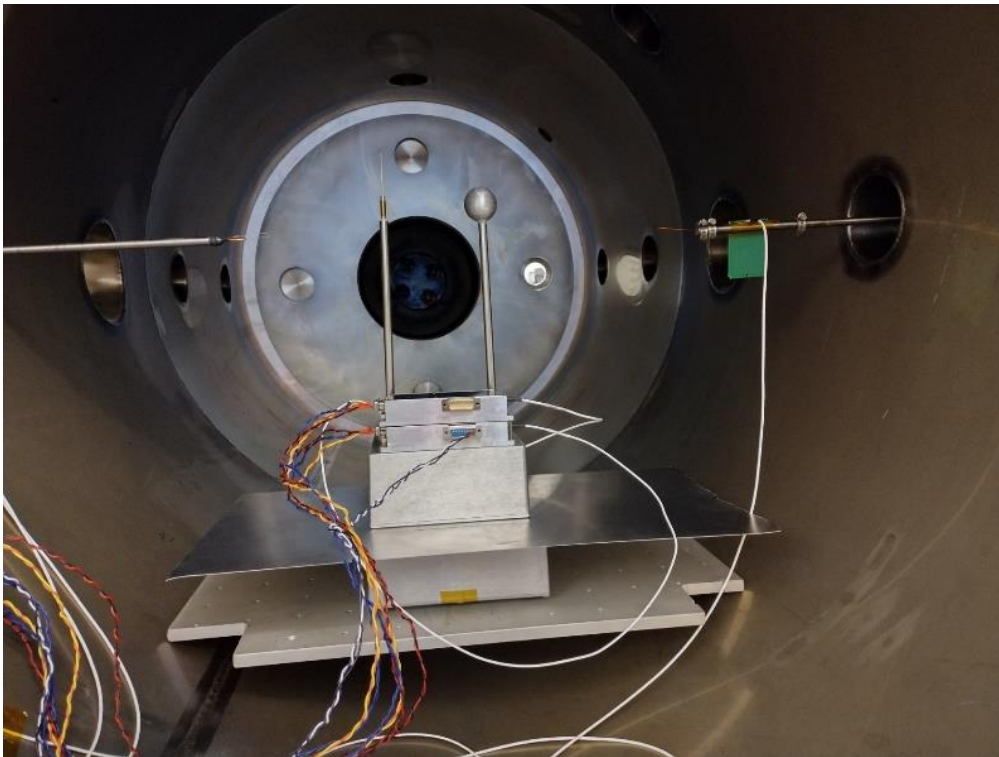
Figure 1. Ulysses magnetometer data from 17 July 2002, around 08:15 UT. (a) Field components as a function of time expressed in RTN coordinates. (b) Field components as a function of ξ , the scaled and rotated x , after the following transformations have been applied: (i) rotation using the minimum variance transformation, (ii) scaling to the asymptotic value of B_x , and (iii) rotation about the new x direction to eliminate the asymptotic value of b_z . The eigenvalues that resulted from the minimum variance analysis [Sonnerup and Scheible, 2000] were 33.88, 15.75, and 1, scaled to the minimum eigenvalue. The minimum variance direction, x , makes an angle of $\theta \approx 32.8^\circ$ with respect to the background magnetic field.

Wheeler, H. R., IV, M. A. Reynolds, and R. L. Hamilton (2015), J. Geophys. Res. Space Physics, doi:10.1002/2014JA020770.

Electric Field

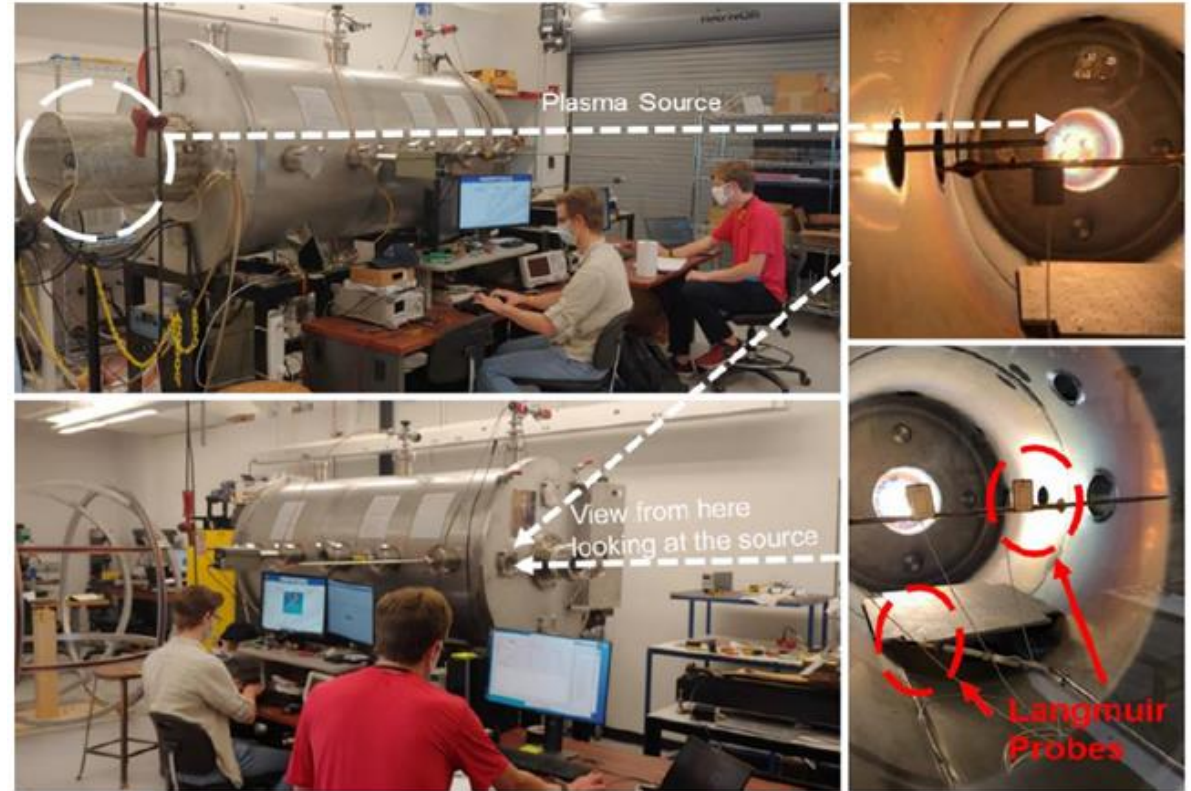


- SAIL Floating probes have 10s of μV noise floor... $<0.1\text{mV/m}$ resolution
- Tested in Plasma Chamber and multiple upcoming rocket flights



Testing in a plasma Chamber !?!

- Concerns with solitons:
 - Strong geometrical attenuation in 3D, as opposed to 1D or 2D
 - Effect on solitons due to density gradients?
 - Effect on particles due to energy lost by the waves *?
- With a LEO plasma source that generates flowing plasma we can test solitons in a plasma chamber with many possibilities
 - Multiple geometries and arrival angles
 - Multiple debris sizes and flow velocities



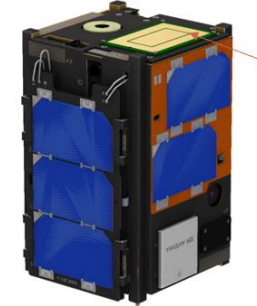
Plasma Chamber setup at ERAU.

MMOD < 1mm : Observations by impact on spacecraft

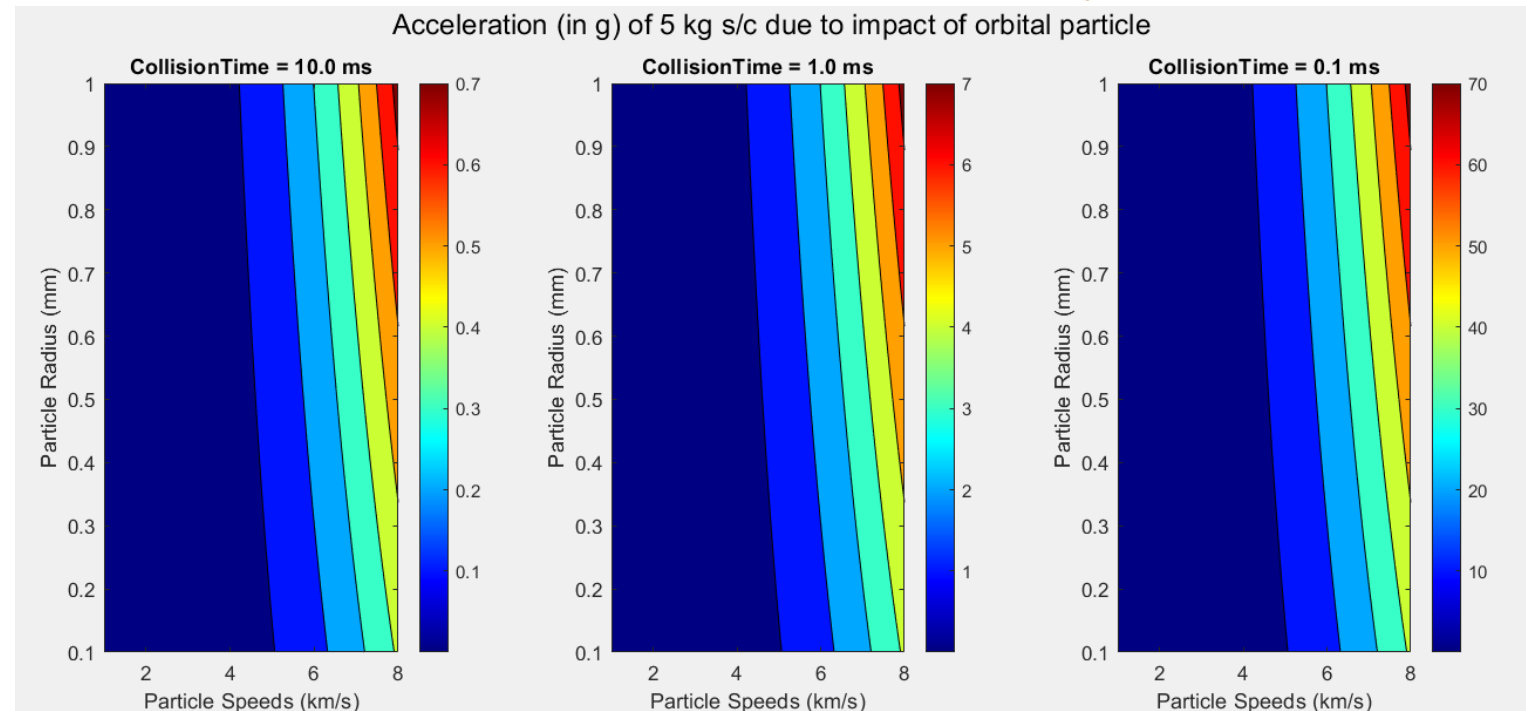


- Oscillations in potential will be seen by Floating Potential Probe
- Changes in local plasma density will be seen by patch Langmuir probes. To right is an image of a 1.5U CubeSat that has a gold patch.
- Accelerations seen by sensitive high sample rate accelerometers

LLITED 1.5U CubeSat



- Plot to the right shows particles coming in varying relative speeds (x-axis), and particle size (y-axis) assuming aluminum spheres. CubeSat mass assumed is 5kg (3U).
- Depending on how long it takes for the particle to come to rest, the acceleration experienced by the CubeSat is shown in 3 columns.
- Takeaway: we need accelerometers sampling at several KHz to catch the impact



Summary



- **SAIL Langmuir probes are low SWaP.**
 - They can conform to any reasonable Length x Width PCB dependent on small satellite bus constraints.
 - Their performance characteristics make them suitable for high cadence density measurements allowing observations of solitons.
 - Deployment of multiple sensors on same spacecraft with optimal distances could potentially enable directional information
 - Plasma chambers with flowing LEO source could be the best first step towards determining instrumentation for soliton detection.
- **SAIL has additional small accelerometers that may be suitable for detecting extremely small debris hits (<1 mm). Multiple of these accelerometers could be deployed on large surface area satellite.**
- **We also meticulously calibrate a COTS magnetometer enabling ~1nT accuracy and precision measurements at 1 Hz. Lower precision at higher sampling rates.**